

WFIRST-AFTA Coronagraphic Instrument Science Yield Modeling Updates

206.01

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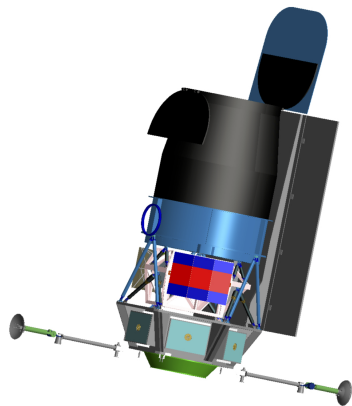


Figure: WFIRST-AFTA
observatory

- 2.4 m aperture primary mirror
- Wide Field Imager/Spectrometer & Integral Field Unit
- Internal Coronagraph with Integral Field Spectrometer
- Overall Dry Mass: 4059 kg
- Structure: high stiffness composites; modular packaging for avionics
- GN&C/Propulsion: inertial pointing, 3-axis stabilized, mono-prop system for stationkeeping & momentum unloading
- Continuous 600 Mbps Ka-band to dedicated ground station
- L2 orbit
- Delta IV Heavy launch

WFIRST Coronagraphic Instruments

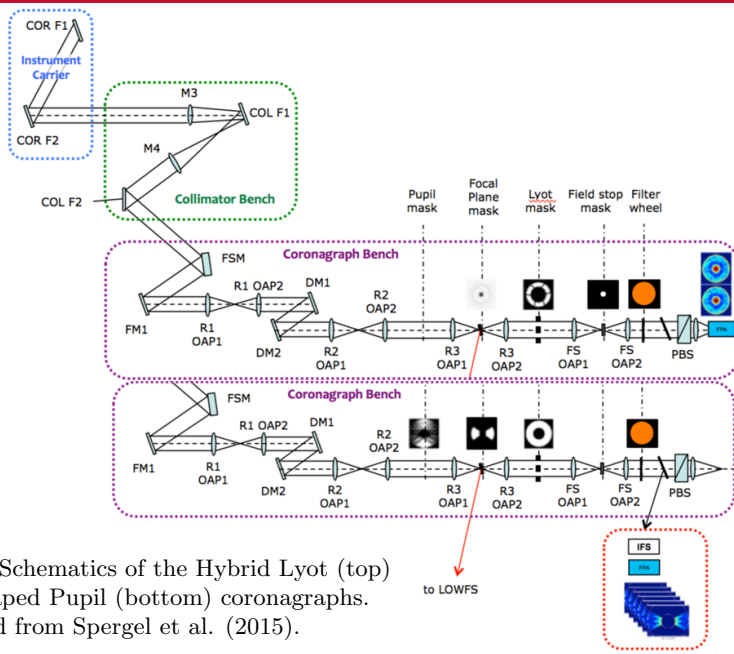
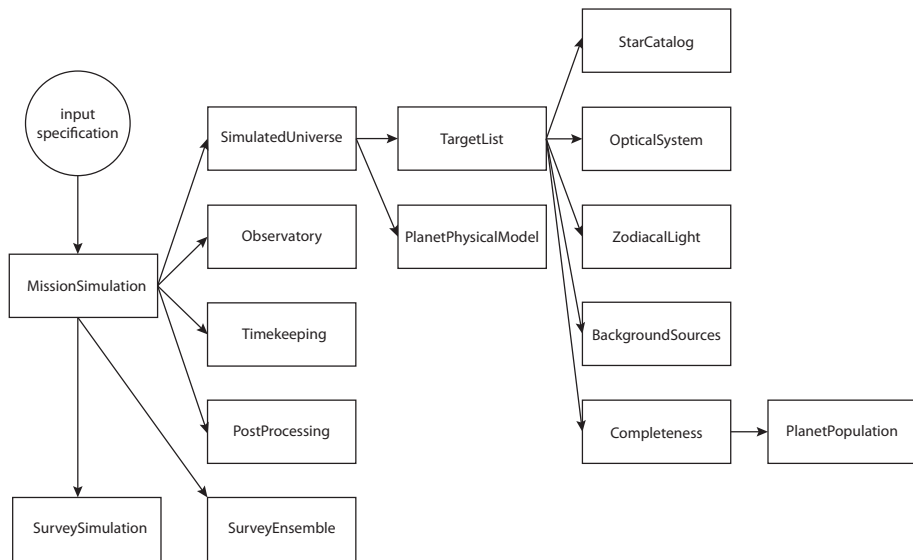


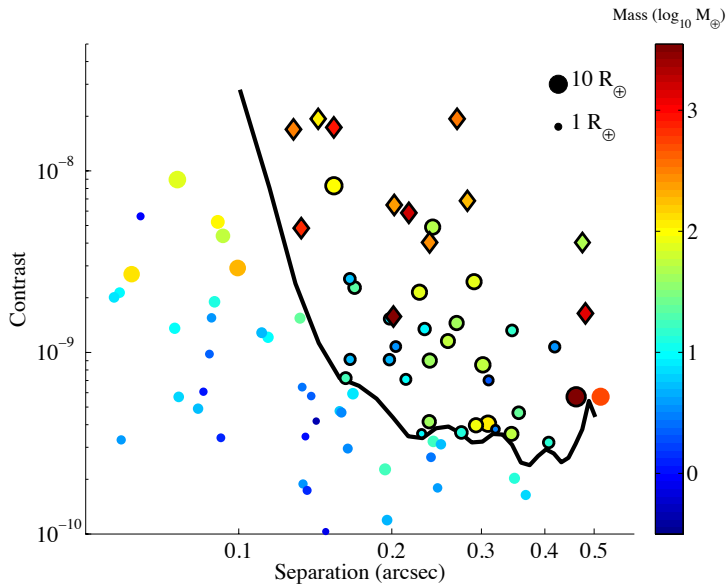
Figure: Schematics of the Hybrid Lyot (top) and Shaped Pupil (bottom) coronagraphs. Adapted from Spergel et al. (2015).



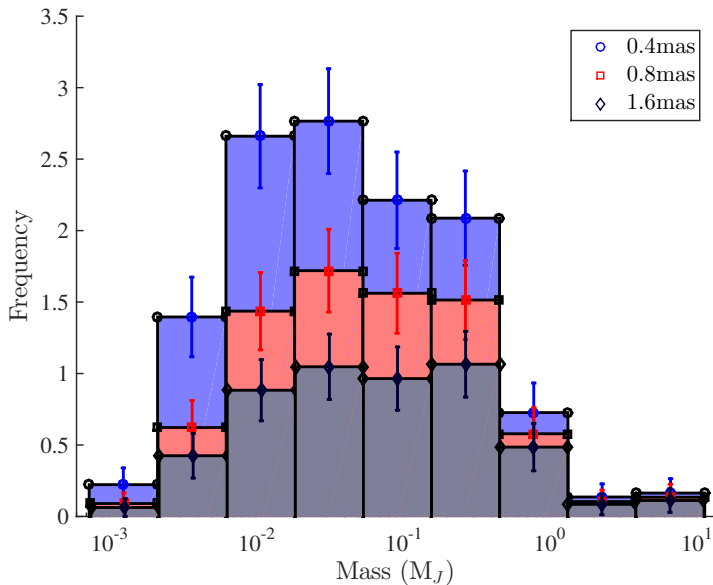
- Describe the range of potential science yields of the WFIRST-AFTA coronagraph and determine:
 - ① What is the optimal proportion of coronagraph time that should be devoted to searching for new planets versus attempting to image known exoplanets?
 - ② What are the best targets and optimal observation times for potential new detections and followup observations?
 - ③ What are the optimal operating points in terms of detection band and permissible false positive rates as a function of angular separation and how does this affect the required integration time for each target?
- Develop and release an open source toolkit for the modeling of space-based planet finders



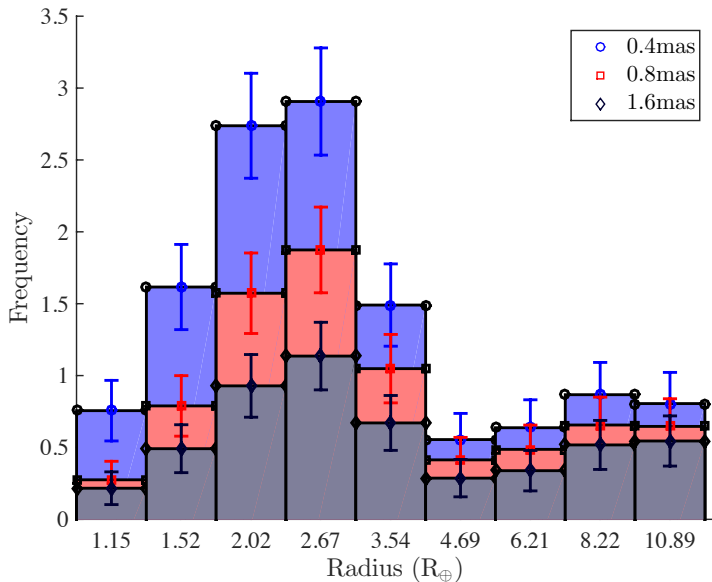
Single Simulation Results

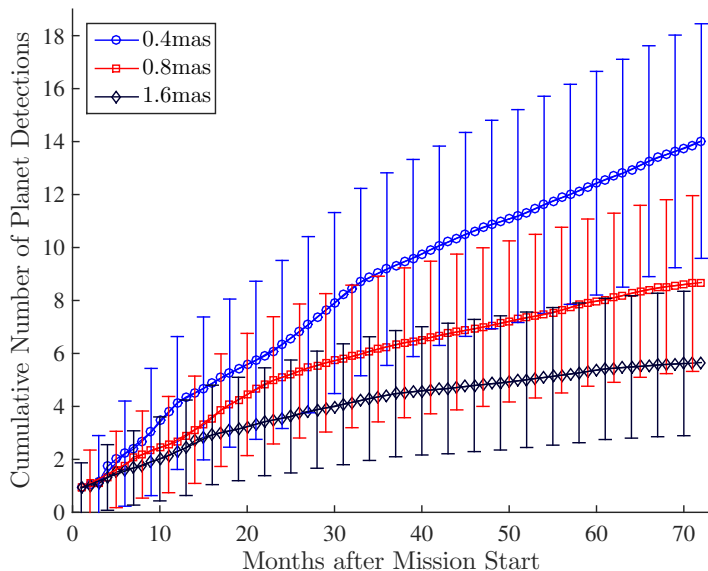


Simulation Ensemble Results



Simulation Ensemble Results







- WFIRST modeling is ongoing - more results soon
- Can (and will) incorporate all of the fantastic results from the other WPS projects and (soon) the SITs
- EXOSIMS is under active development
 - Please see <https://github.com/dsavransky/EXOSIMS> - in particular the ICD and as-built documentation
 - Comments and pull requests very very welcome
- Please see Morgan et al. (305.01)